

Preliminary Results from Dwarfs4MOSAIC:

A two-dimensional study of low-mass, high star-formation galaxies at low redshift

PhD thesis supervised by:
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María Chillarón Víctor

IV Reunión Científica de GUAIX
December 17th 2025



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Dwarfs4MOSAIC ITP

La Palma International Time Program



MEGARA @ GTC



WEAVE @ WHT



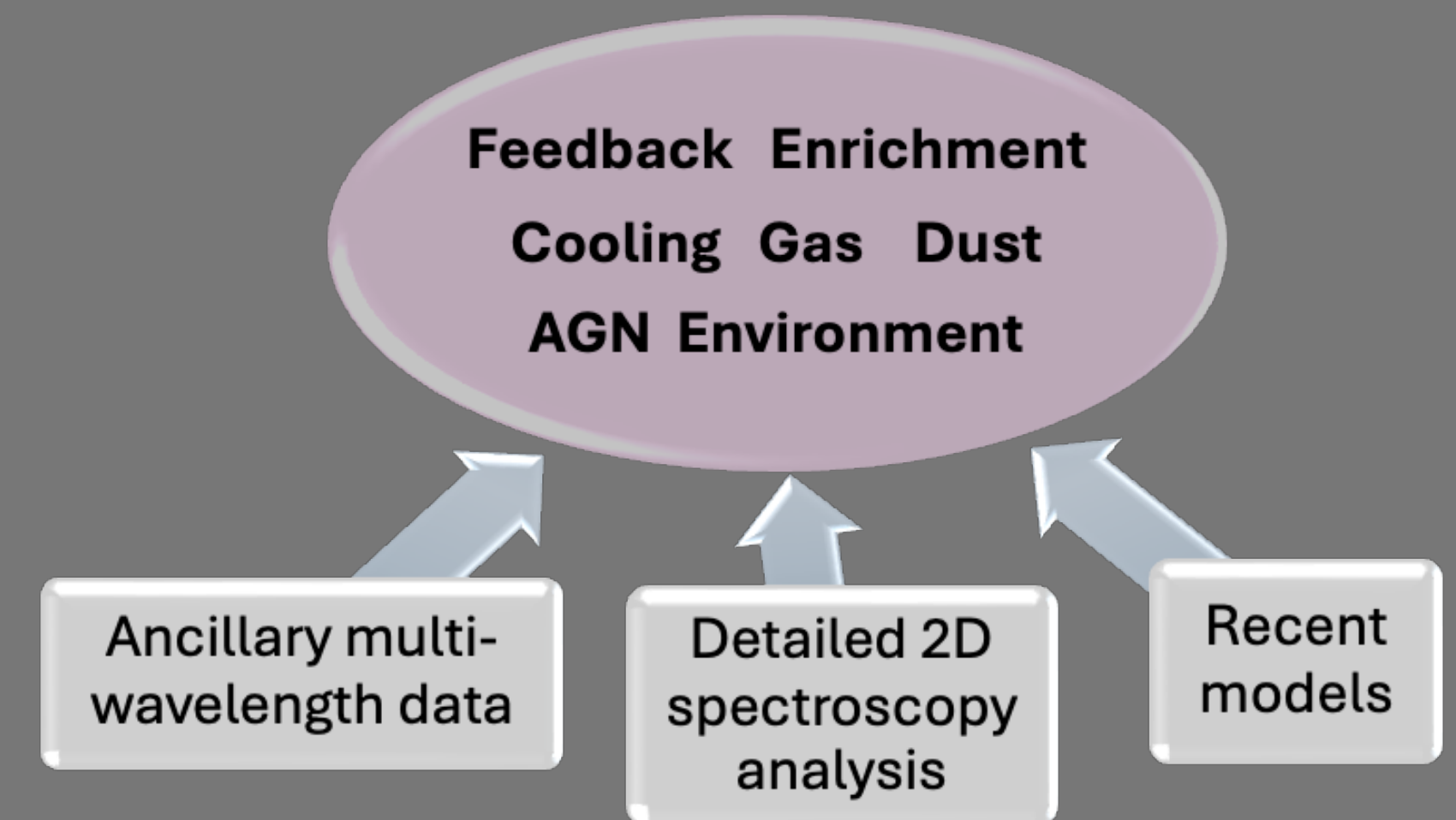
WFC @ INT



10:0 @ LT



A pilot study for MOSAIC at ELT
A 2D study of low-mass star-forming galaxies
as low-redshift analogs to reionization-epoch
primeval galaxies. PI Jesús Gallego



Dwarfs4MOSAIC ITP

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SBS 0335-052W
Mrk1486
J1105+4444

Dwarfs4MOSAIC ITP

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10.4m Gran Telescopio Canarias
Spectrograph based on fibers

* IFU mode

623 spaxels (567 LCB + 56 sky) of 0.62''

VPH Name	Setup	R _{FWHM}	$\lambda_1 - \lambda_2$ (Å)	λ_c (Å)
VP405-LR	LR-U	6028	3653 - 4386	4051
VP480-LR	LR-B	6059	4332 - 5196	4800
VP675-LR	LR-R	6099	6094 - 7300	6747

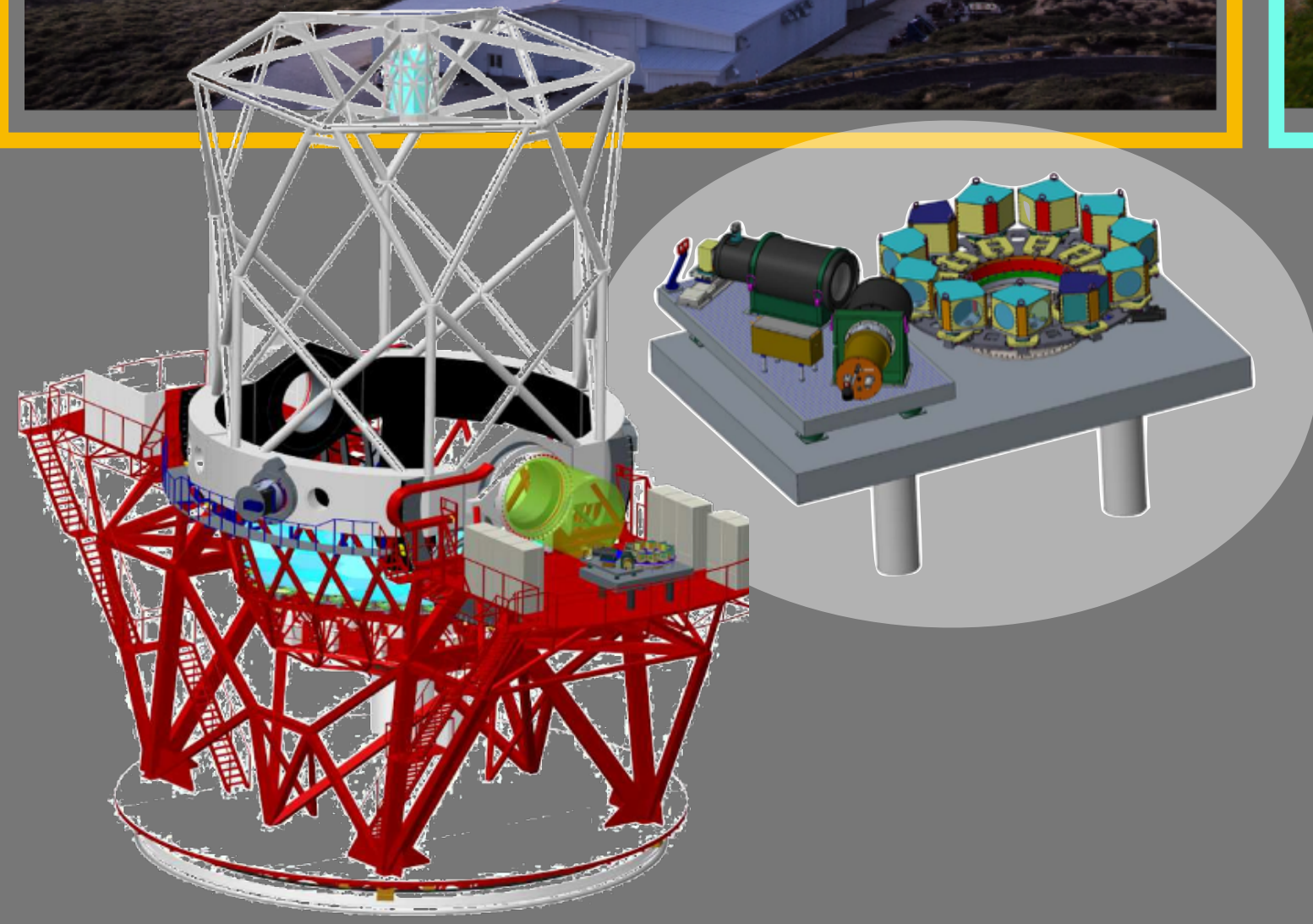
MEGARA instrument and Gran Telescopio Canarias.
Credit: Gran Telescopio Canarias (GTC)

Dwarfs4MOSAIC ITP

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MEGARA @ GTC

WEAVE @ WHT



MEGARA instrument and Gran Telescopio Canarias.
Credit: Gran Telescopio Canarias (GTC)



WEAVE on the WHT. Credit Javier Méndez

SBS 0335-052W
Mrk1486
J1105+4444

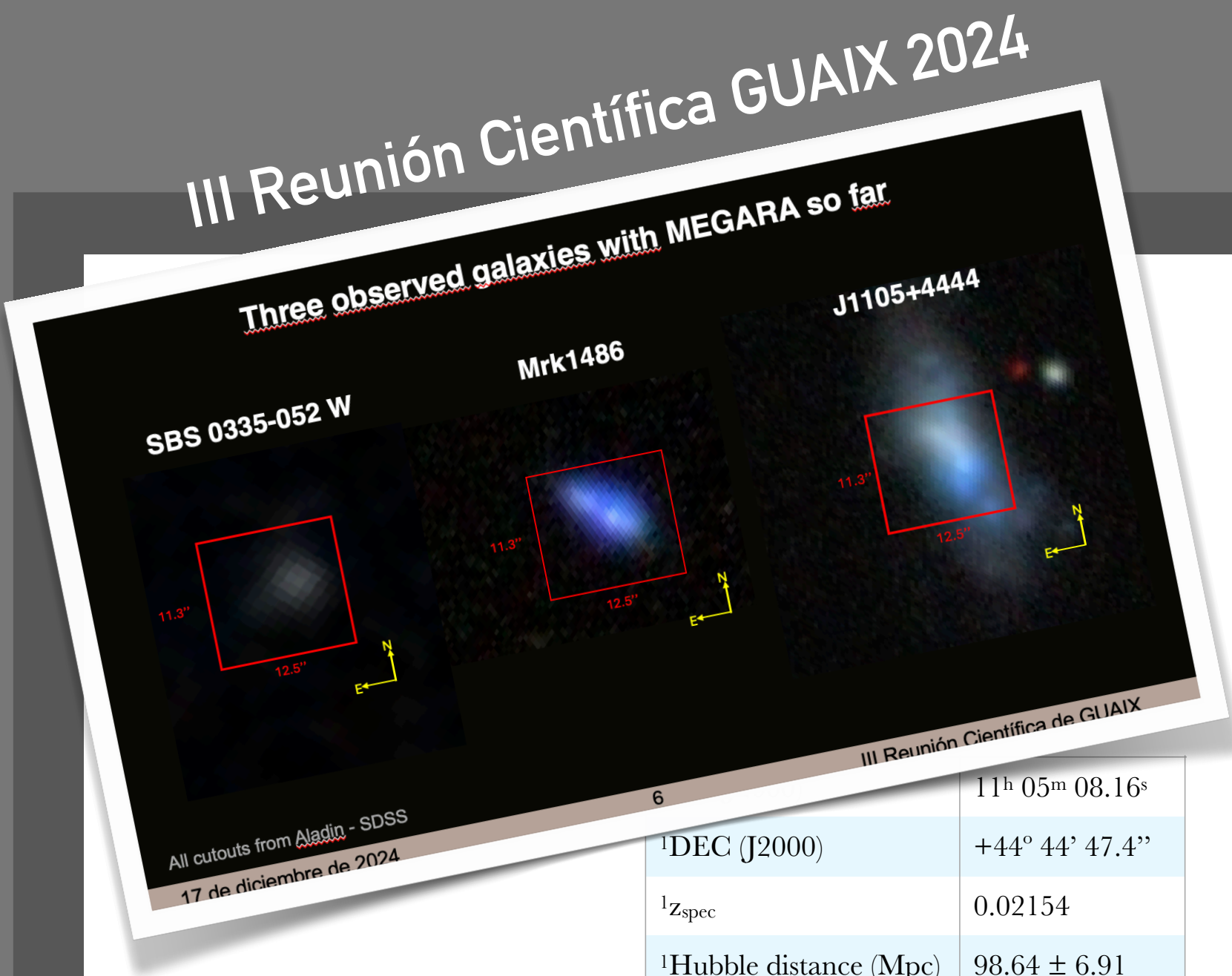
M81 DwarfB
UGC4483
UGC9128
NGC4163
UGC6456
VCC0545
VCC1348
VCC1389

4.2m William Herschel Telescope
* LIFU (large integral-field unit)
603 fibers (547 + 56 sky) of 2.6''
Low-resolution observations:
- Blue arm (3660-6060 Å)
- Red arm (5790-9590 Å)

Reduction pipeline: CASU Cambridge

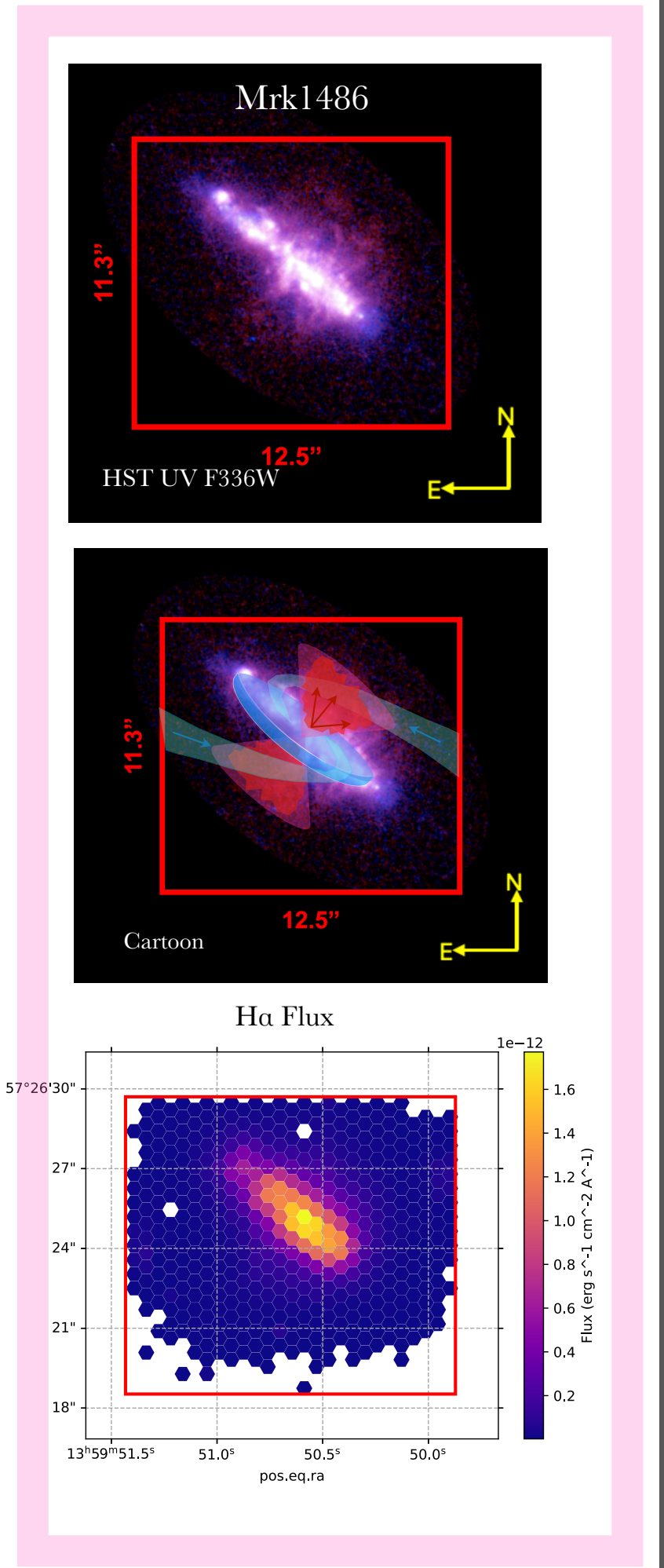
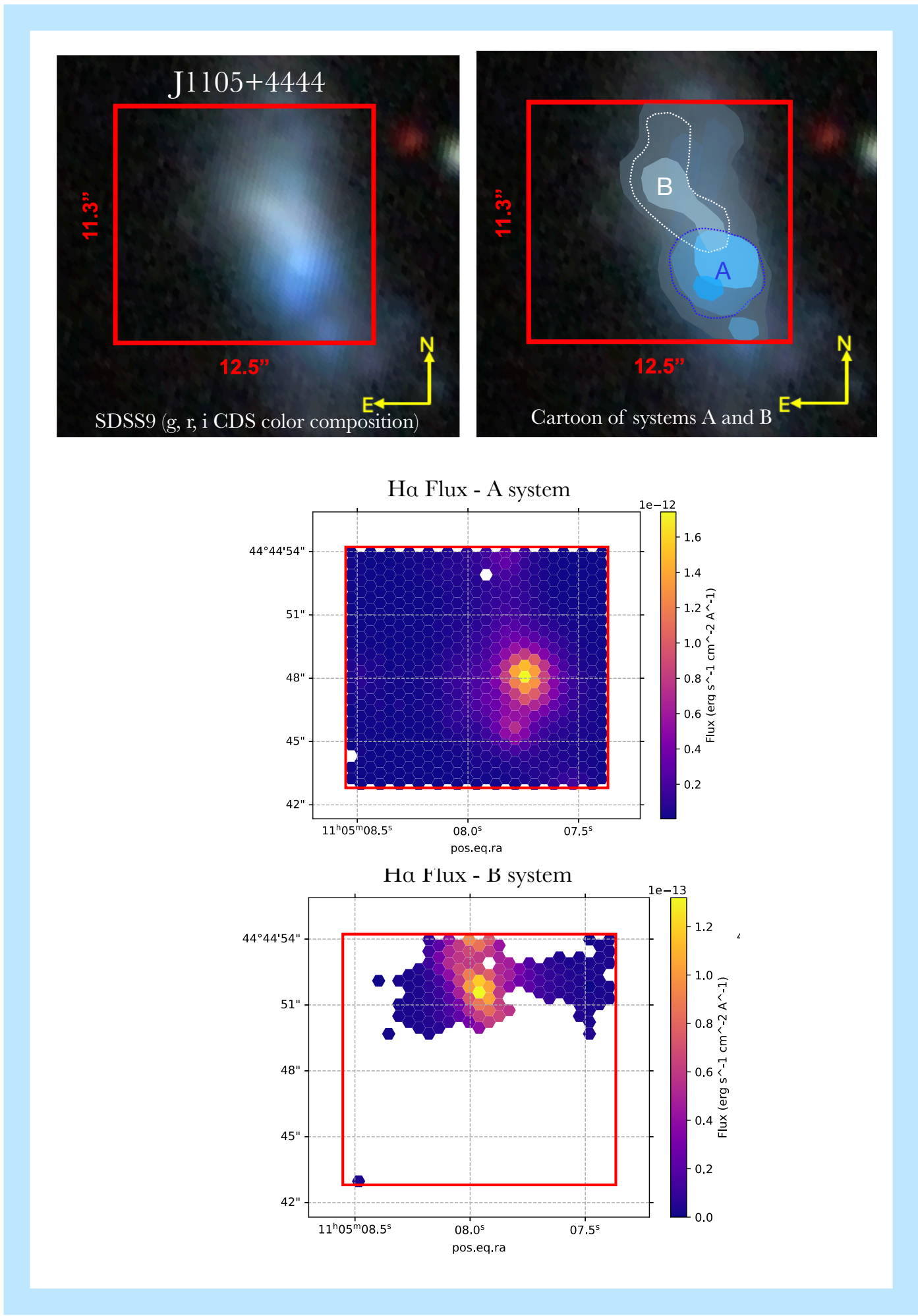
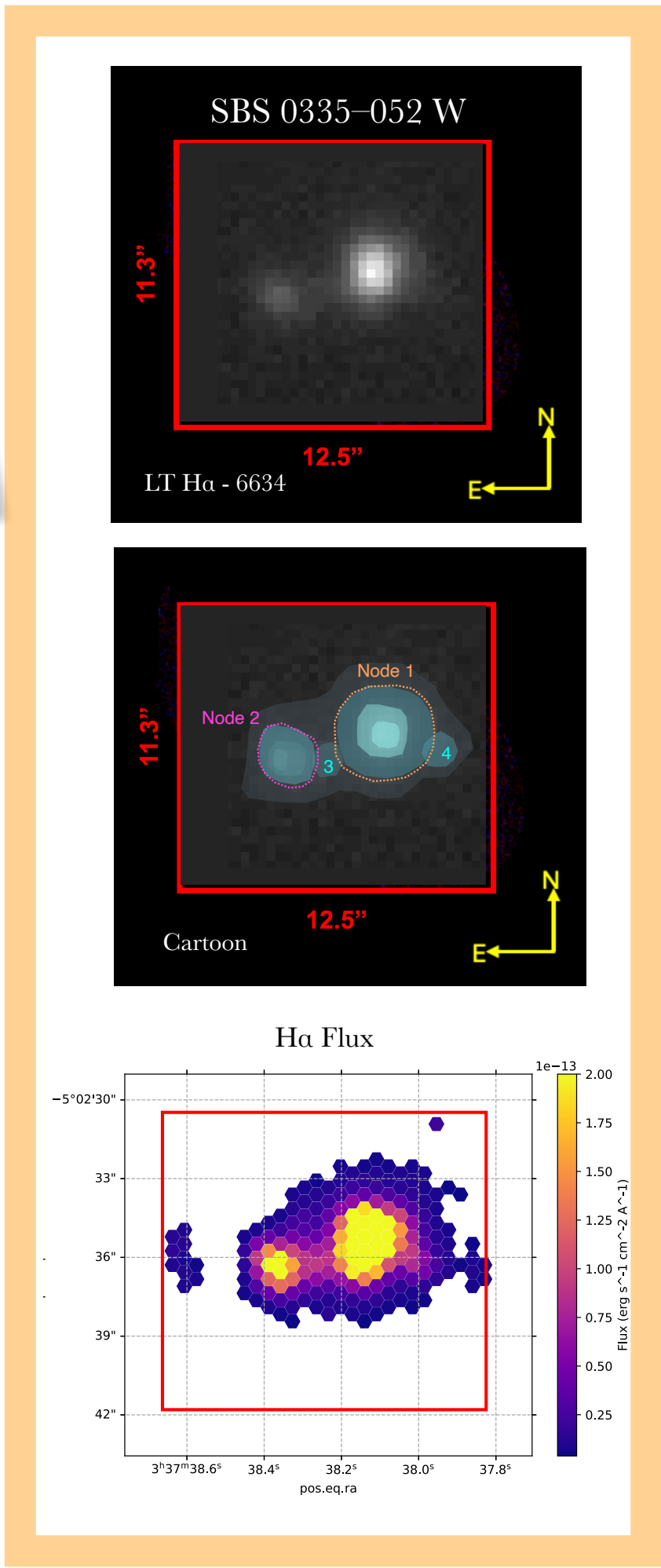
Some of the progress made this year

MEGARA galaxies



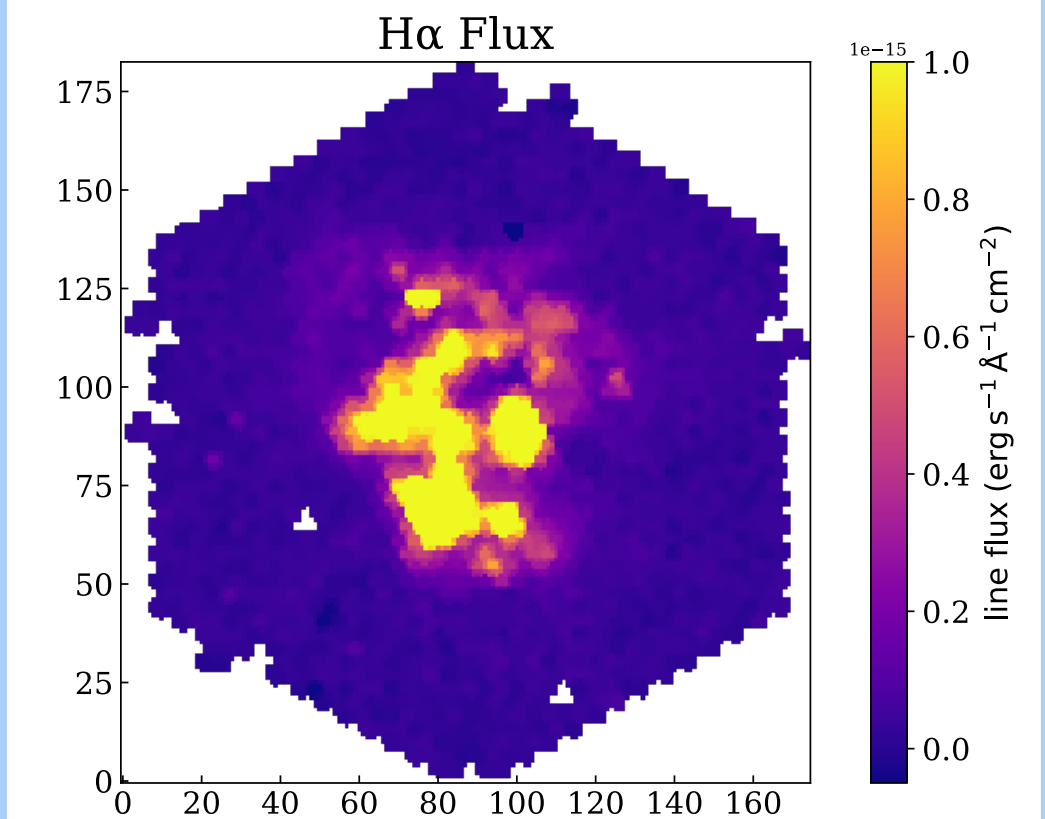
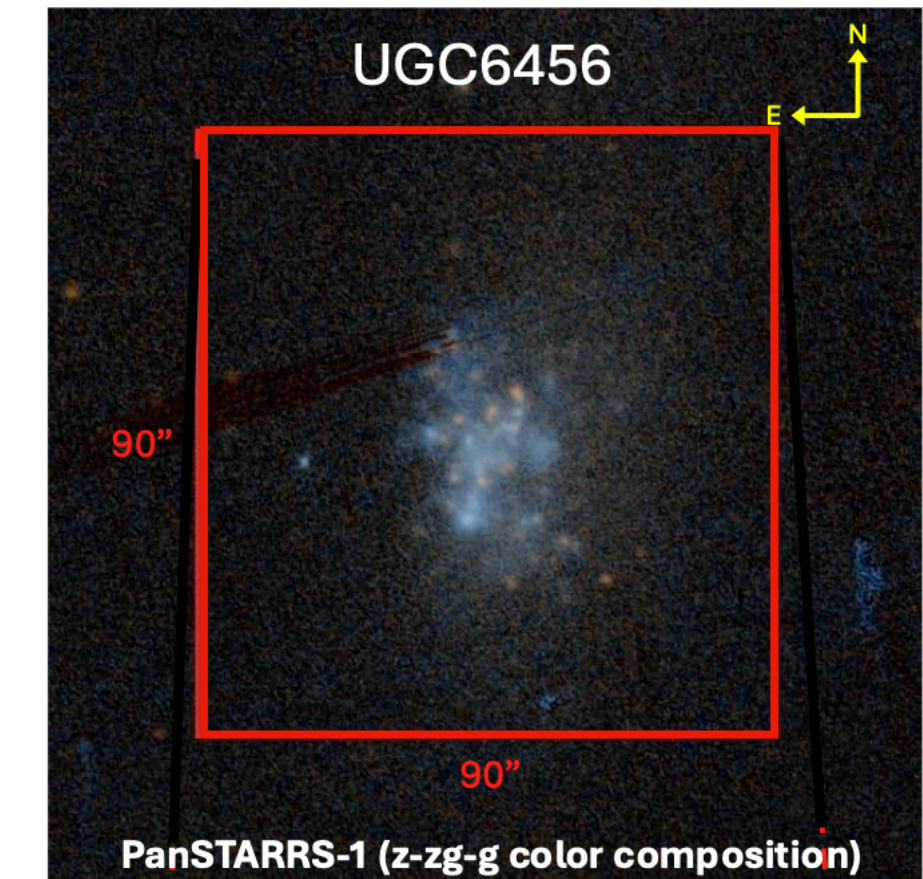
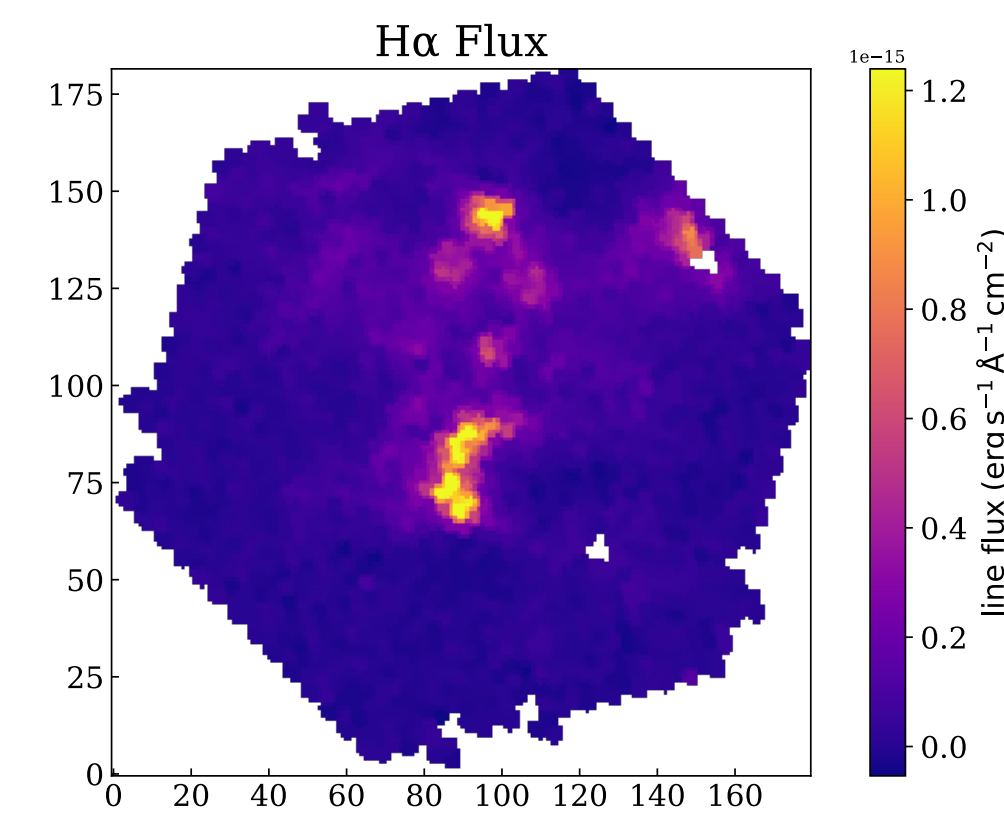
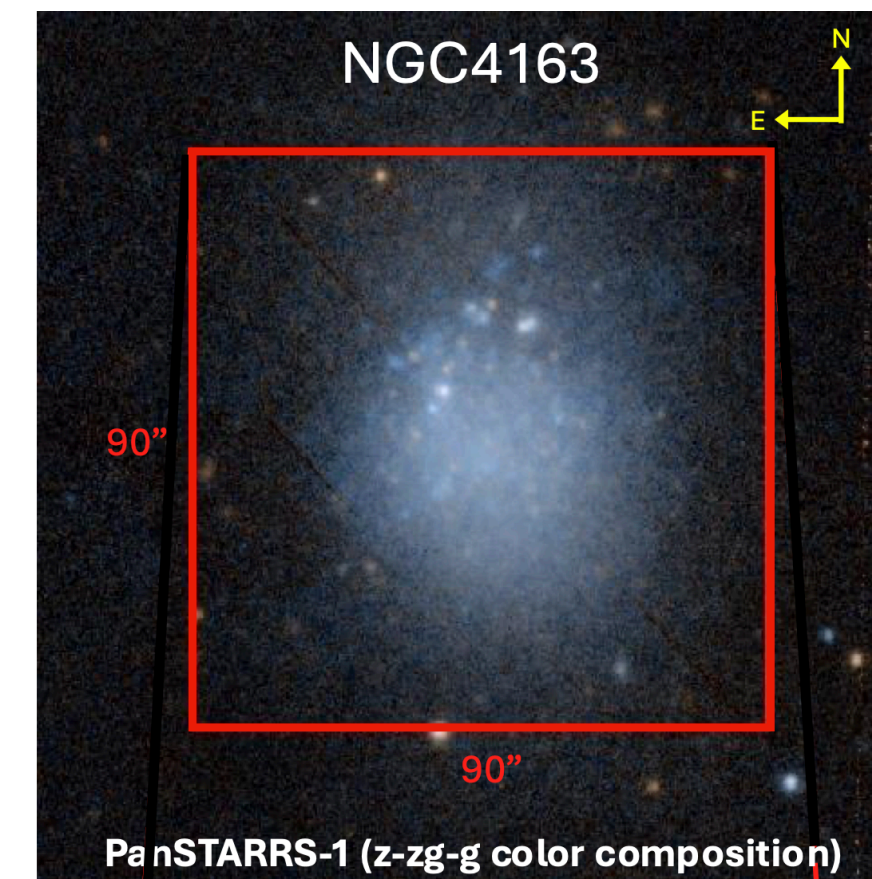
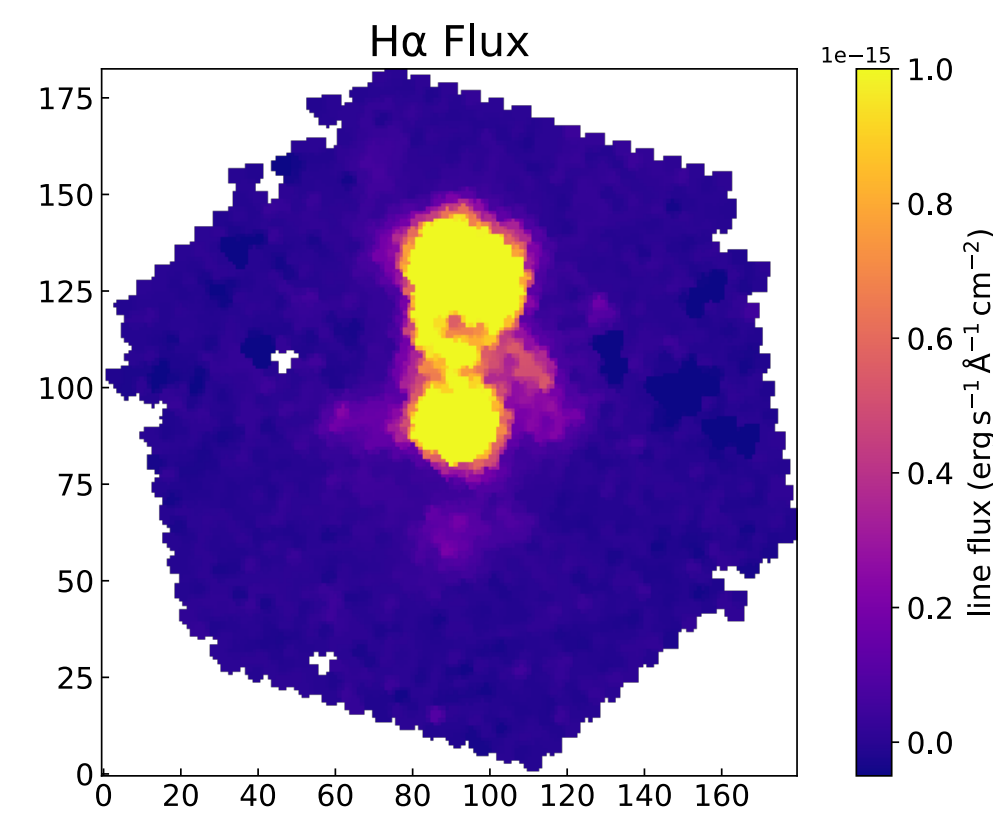
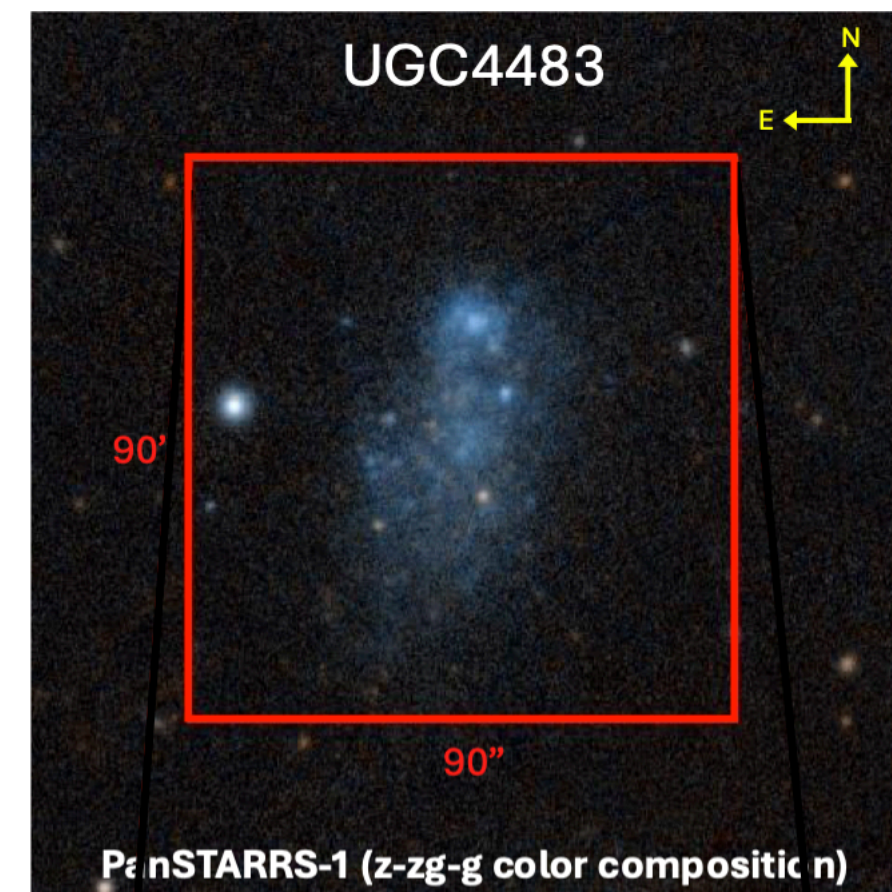
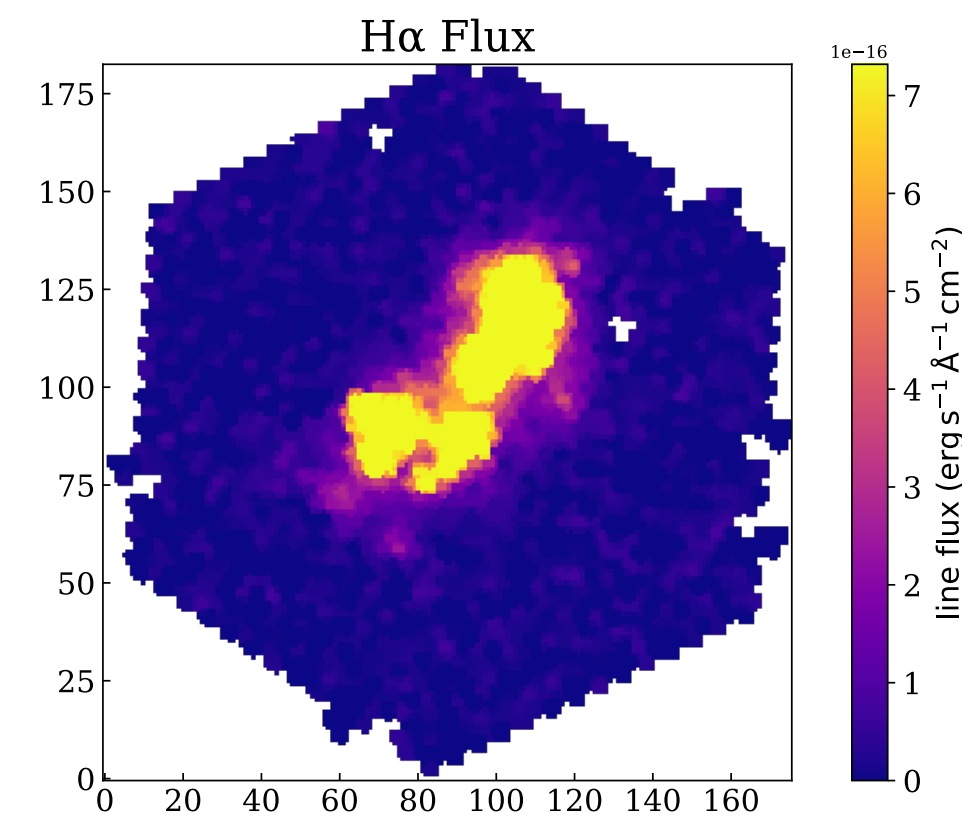
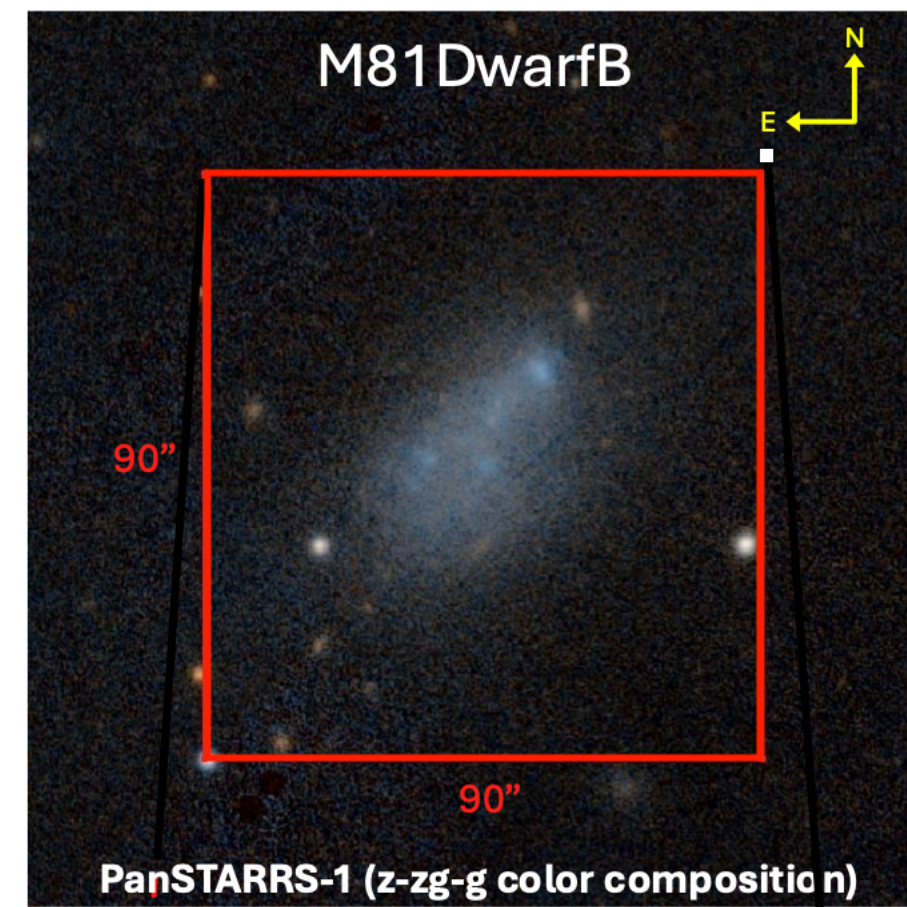
¹ RA (J2000)	03 ^h 37 ^m 44.06 ^s	11 ^h 05 ^m 08.16 ^s
¹ DEC (J2000)	-05° 02' 40.2"	+44° 44' 47.4"
¹ z _{spec}	0.013585	0.02154
^{1,2} Hubble distance (Mpc)	57.11 ± 4.00	98.64 ± 6.91
³ Diameter (arcsec; kpc)	5.2 ; 1.3	40.33 ; 19.79
⁴ 12+log(O/H)	7.22 ± 0.07 (node 1) 7.01 ± 0.07 (node 2)	8.12
⁵ M*/(M*+M _{gas})	0.016	8.69
⁵ M _{dyn} (M _⊙)	7.9 · 10 ⁹	3.63
⁵ M _{HII} (M _⊙)	5.8 · 10 ⁸	

¹ RA (J2000)	13 ^h 59 ^m 50.90 ^s
¹ DEC (J2000)	+57° 26' 22.99"
² z _{spec}	0.033841
¹ Hubble distance (Mpc)	151.20 ± 10.58
¹ Diameter (arcsec; kpc)	13.38 ; 10.08
³ 12+log(O/H)	7.8
² log(M*/M _⊙)	9.3 ± 0.2
⁴ SFR (M _⊙ /yr)	3.6 ± 0.7



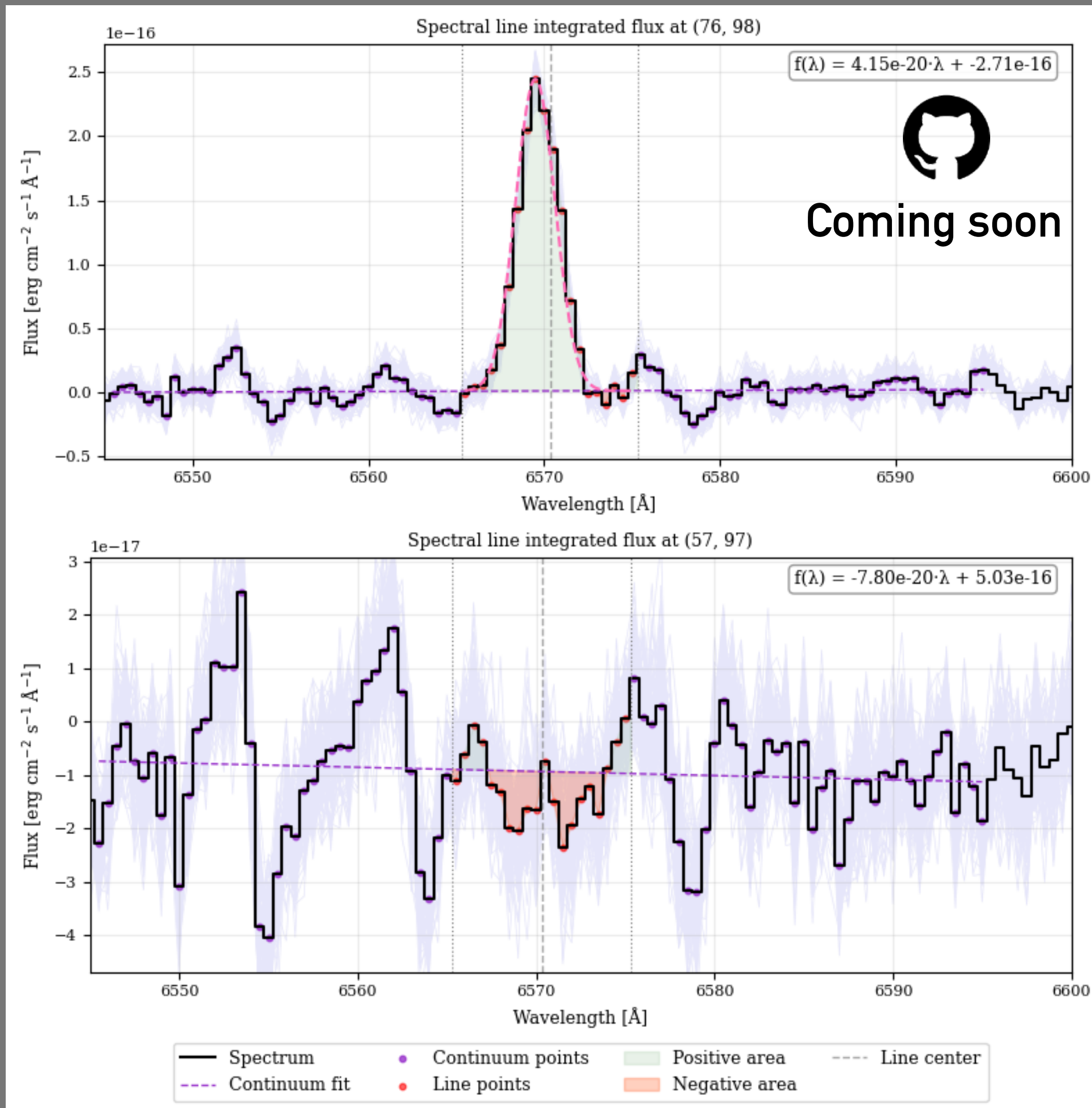
Some of the progress made this year

WEAVE galaxies



Some of the progress made this year

WEAVE galaxies

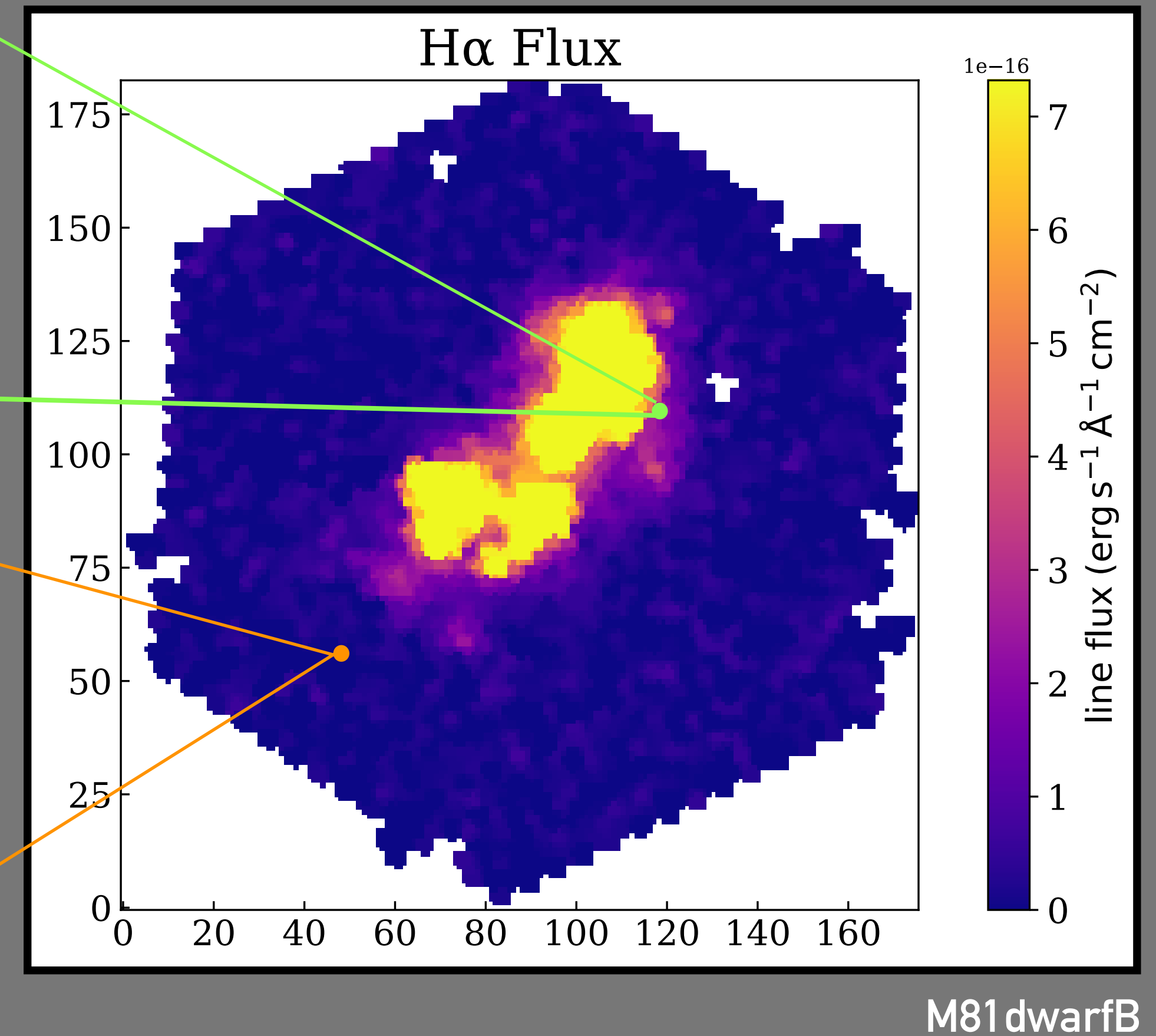
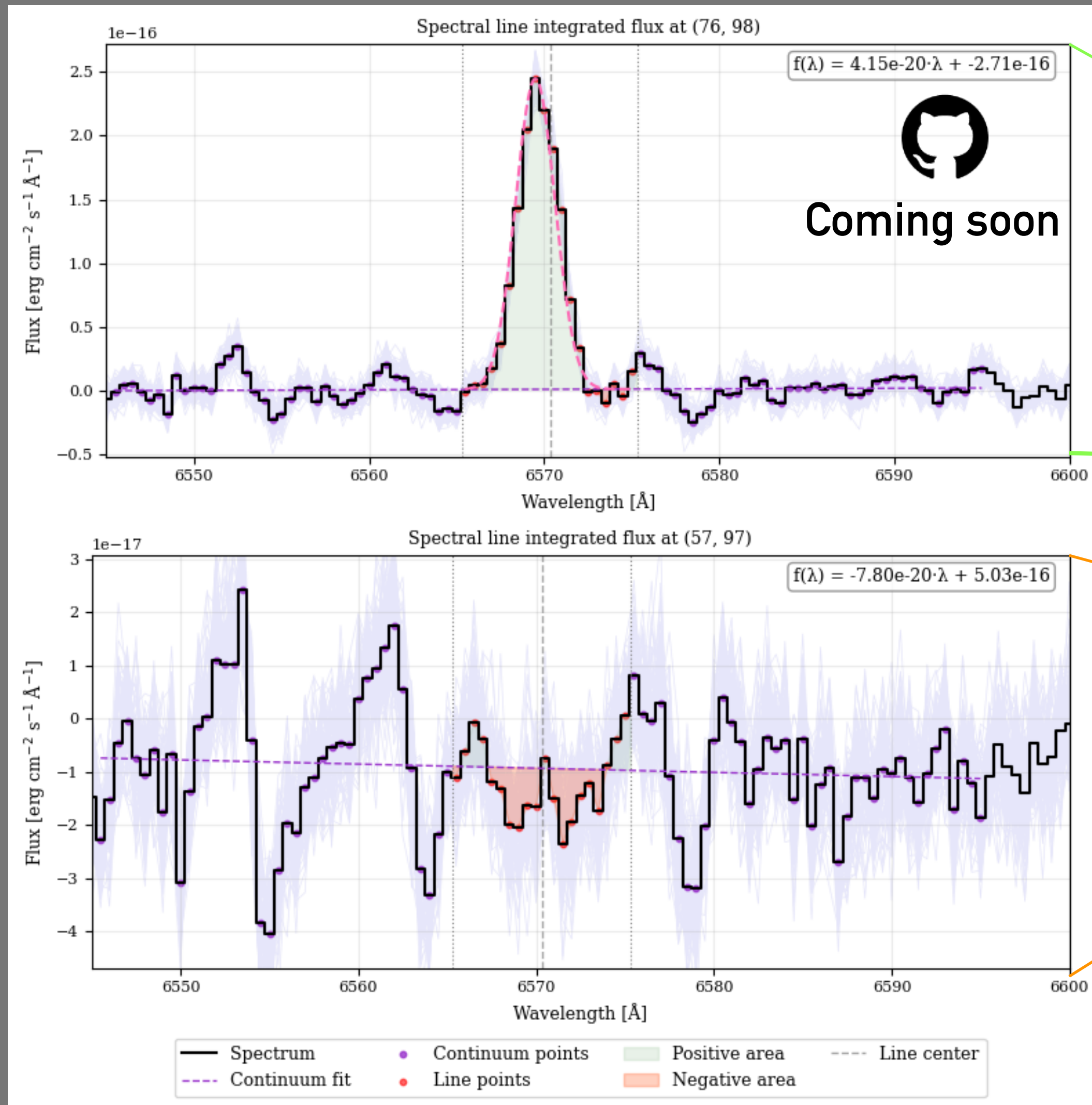


Code for spectral line analysis:

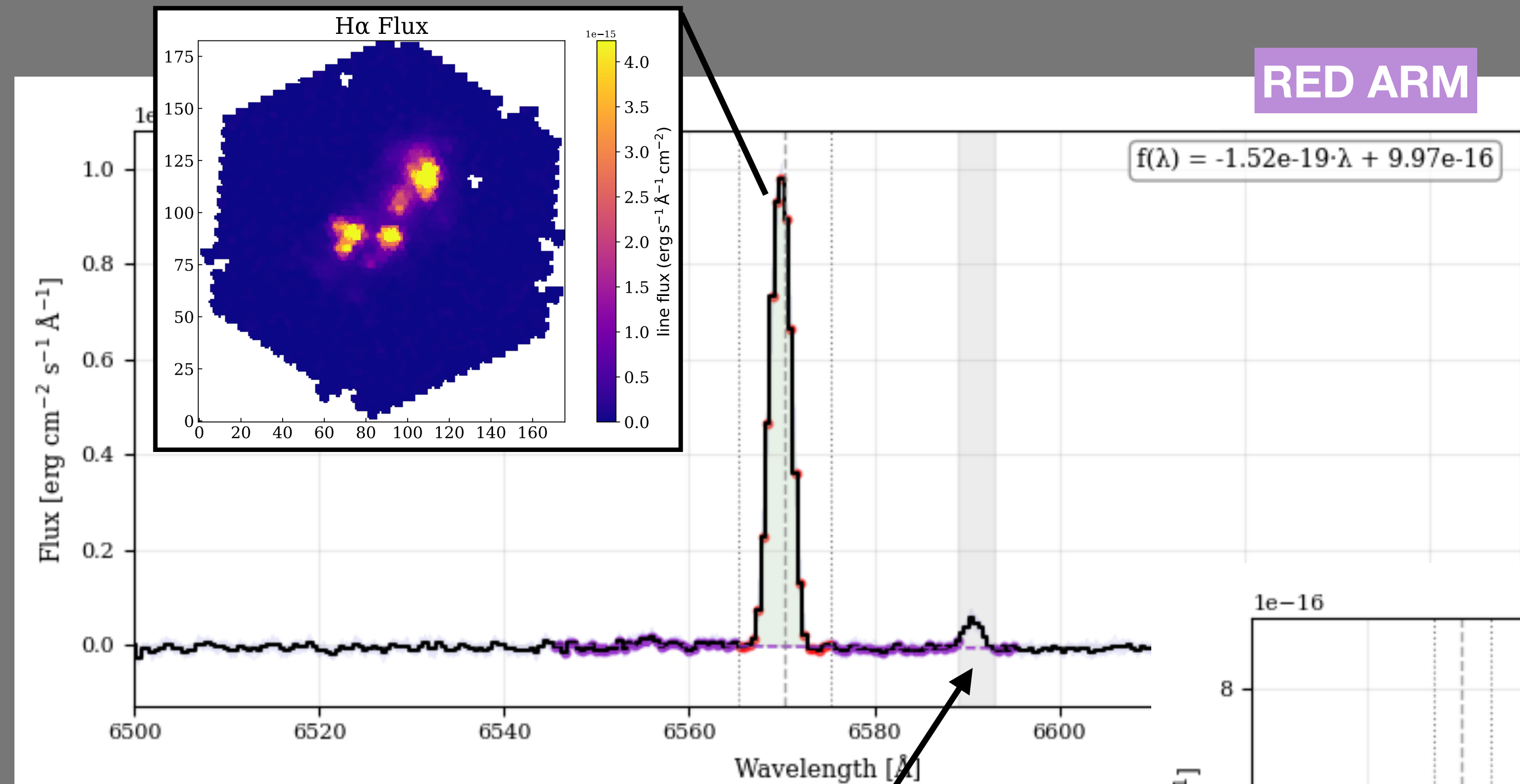
- Line flux measurement using trapezoidal integration
- Gaussian fitting for parameter estimation (σ , amplitude, central wavelength)
- Flux maps, line ratios, and other properties such as velocity or velocity dispersion

Some of the progress made this year

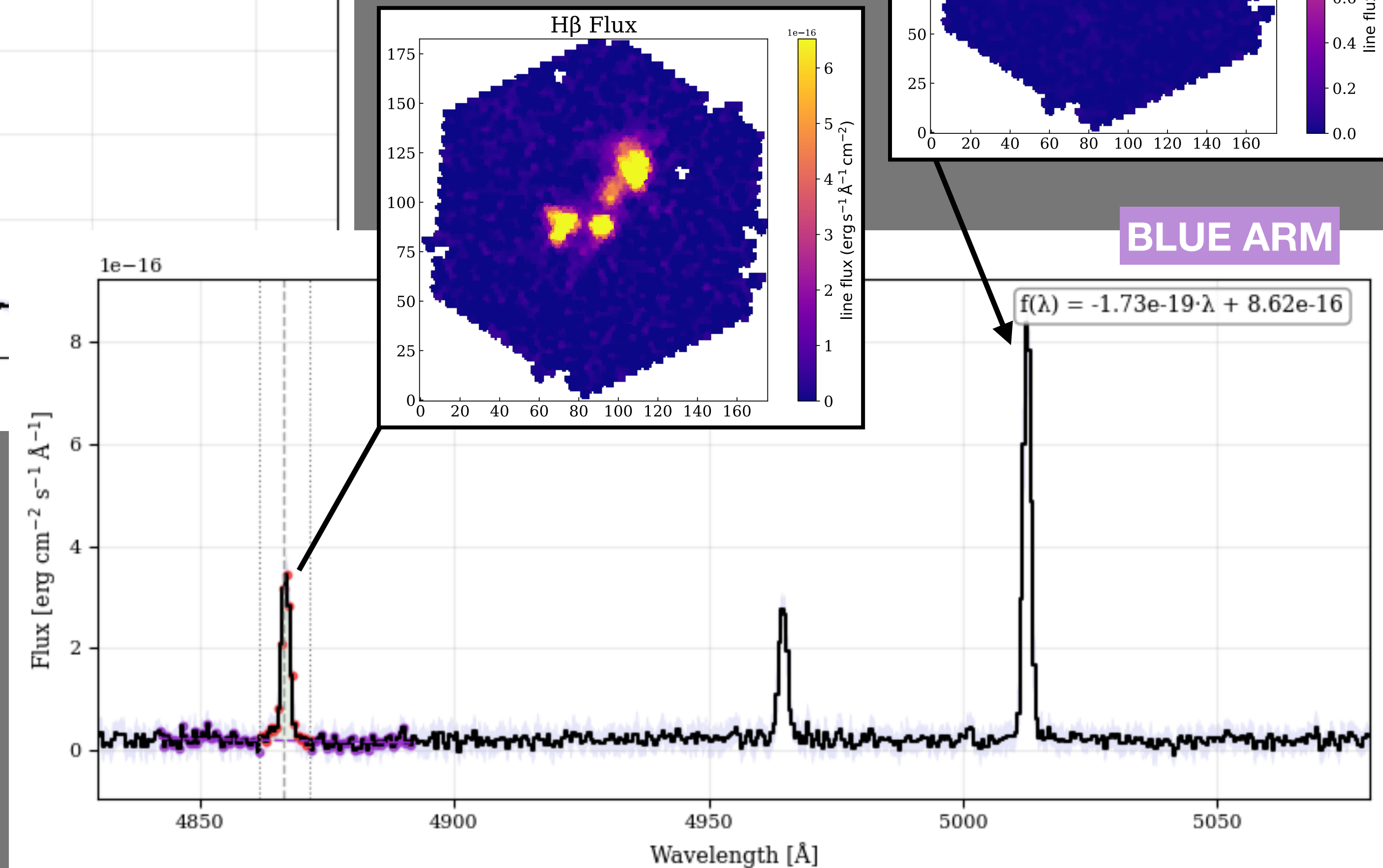
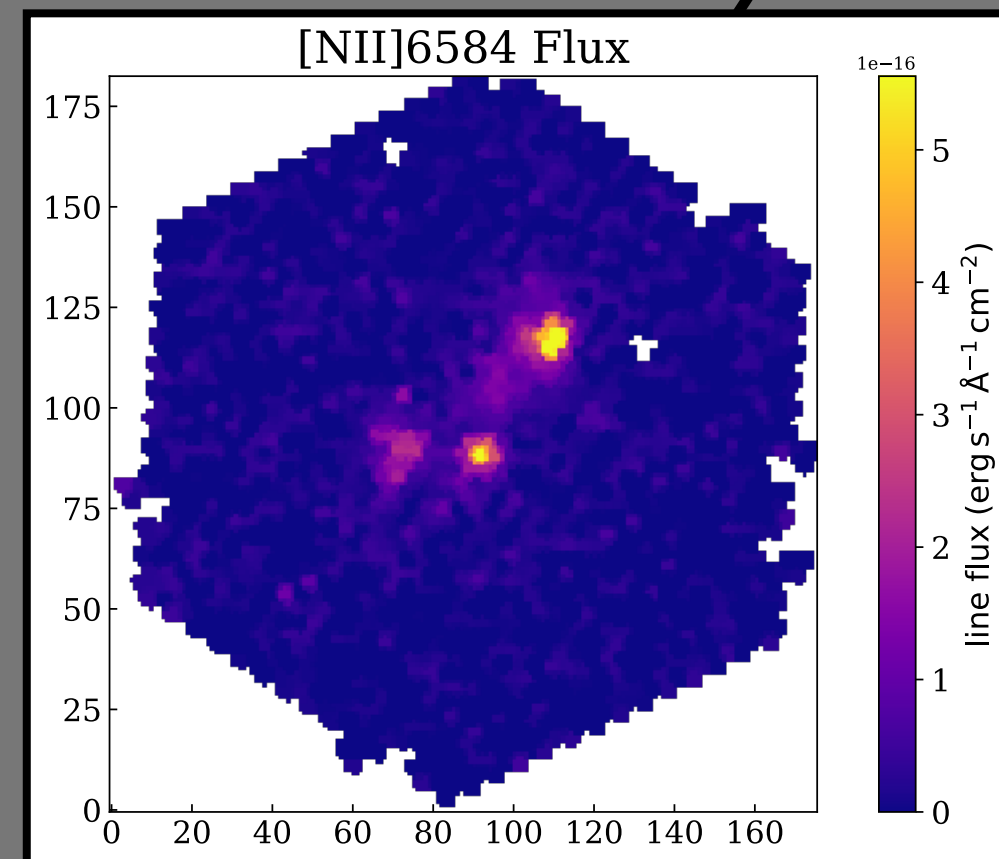
WEAVE galaxies



Some of the progress made this year



M81 dwarfB



Some of the progress made this year

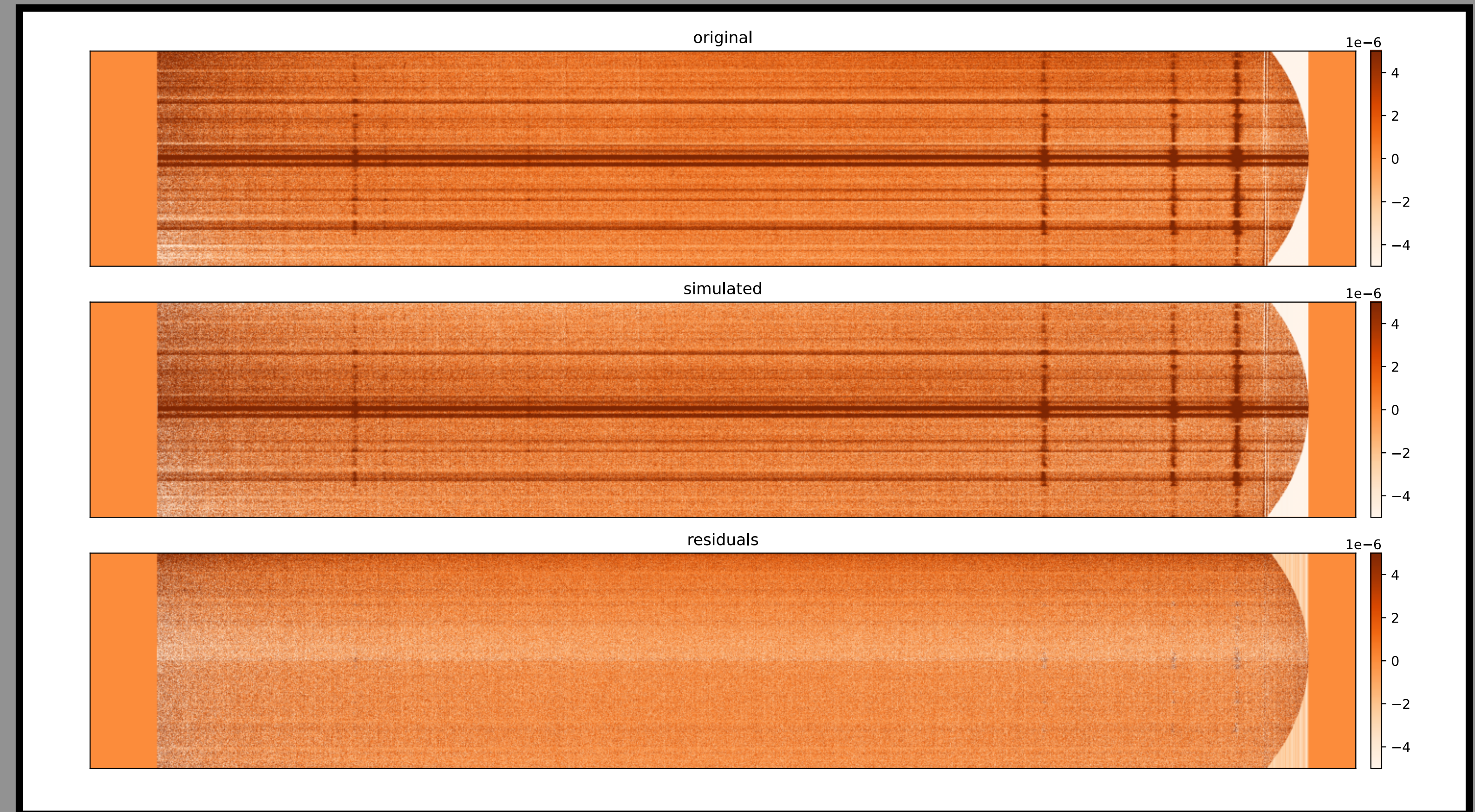
MEGARA Data Reduction Pipeline



megaradrpsimul
mchillaron

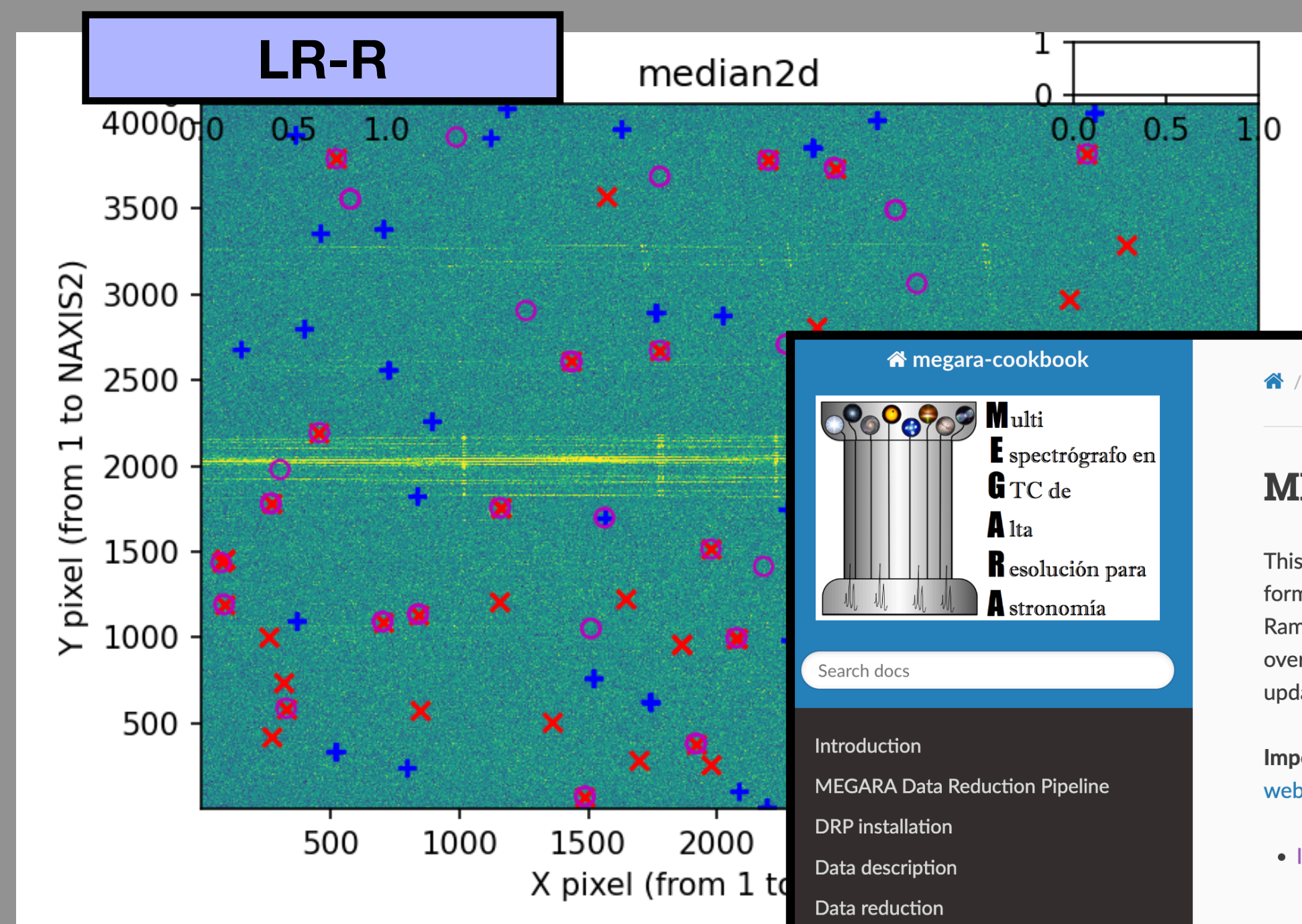
Simulation of raw MEGARA images with Poisson statistics and automated data reduction.

Uncertainty estimation through bootstrapping: error estimation by repeating the analysis multiple times, each time using a new resampled dataset reconstructed from the original data.



Some of the progress made this year

MEGARA Data Reduction Pipeline



crmask
guaix-ucm/numina
nicocardiel

megara-cookbook

Multi
E spectrógrafo en
G TC de
A lta
R esolución para
A stronomía

Search docs

- Introduction
- MEGARA Data Reduction Pipeline
- DRP installation
- Data description
- Data reduction
- Healing Defective Traces
- Problem with bright line tails
- CR not removed by median stacking
- MEGARA Tools
- Known ISSUES
- Acronyms

MEGARA data reduction cookbook

This webpage is an updated version of the original cookbook (dated 07/07/2020), available in PDF format at this [zenodo link](#). That document was prepared by África Castillo Morales, Sergio Pascual Ramírez and Armando Gil de Paz. As the aforementioned document has become slightly outdated over time, we have decided to move the documentation to this website that allows for easier updates. Nicolás Cardiel López and Mario Chamorro Cazorla have also joined this task.

Important: We strongly recommend using the most up-to-date information available on this [webpage](#).

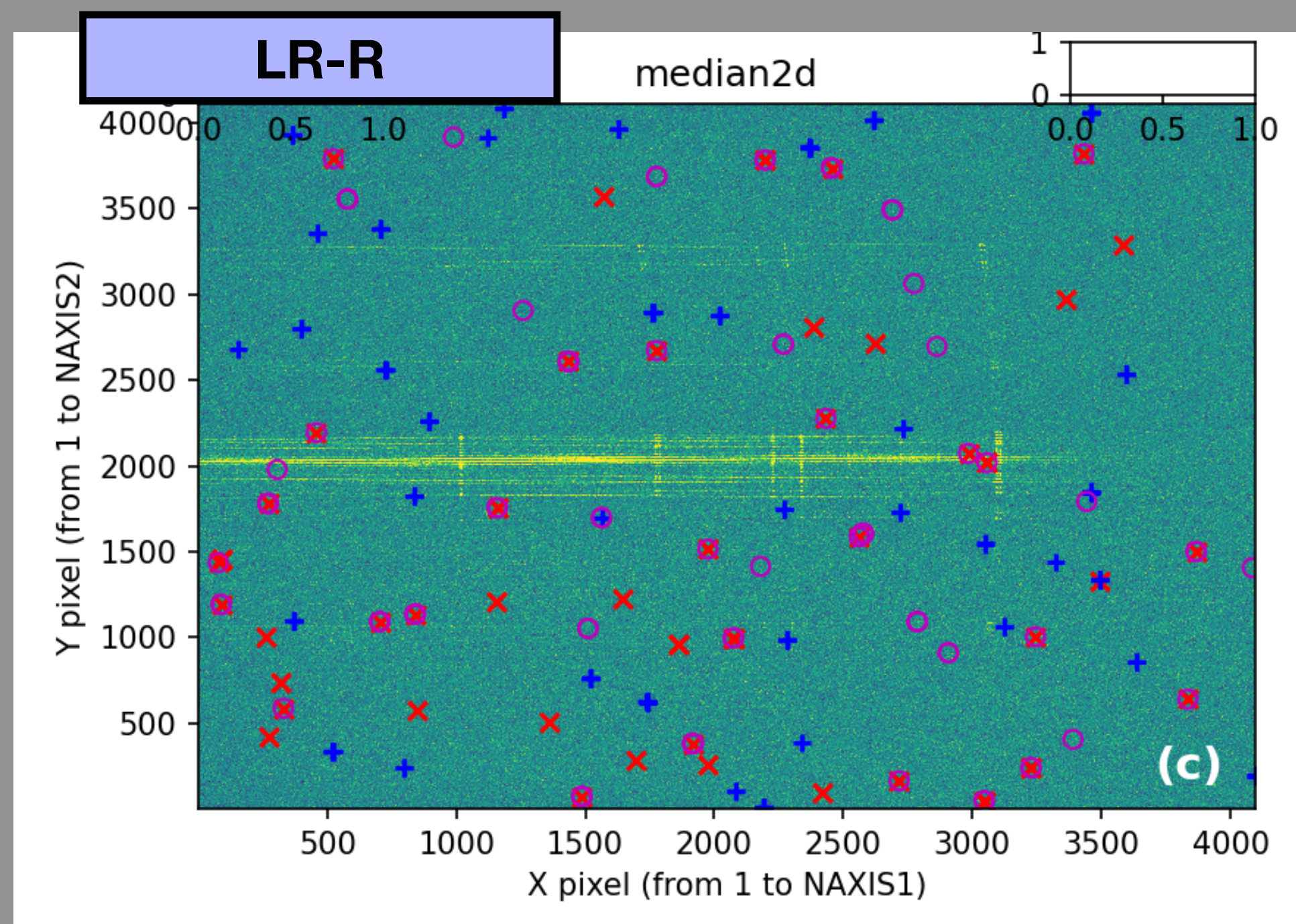
- [Introduction](#)
 - [Scope](#)
 - [MEGARA instrument](#)
- [MEGARA Data Reduction Pipeline](#)
- [DRP installation](#)
 - [Install with pip](#)
 - [Install in conda](#)
 - [Development version](#)
- [Data description](#)
 - [Raw Data](#)
 - [Pipeline Products](#)
- [Data reduction](#)

Median combination effectively cleans pixels affected by cosmic rays when they are present in only a single exposure.

Pixels affected by cosmic rays in more than one exposure (double or triple CRs) have been found to remain uncorrected in the final image.

Some of the progress made this year

MEGARA Data Reduction Pipeline



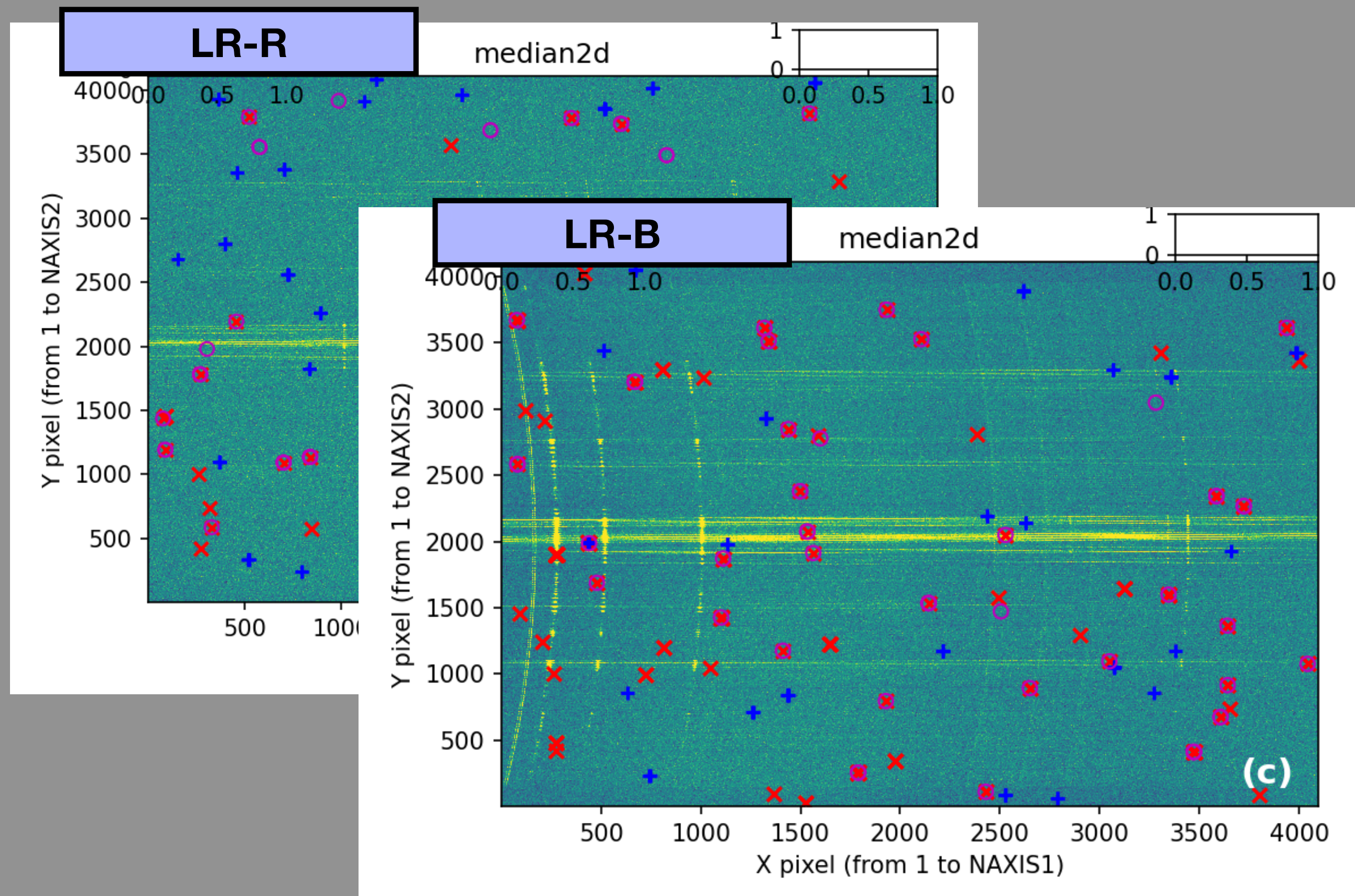
crmask
guaix-ucm/numina
nicocardiel

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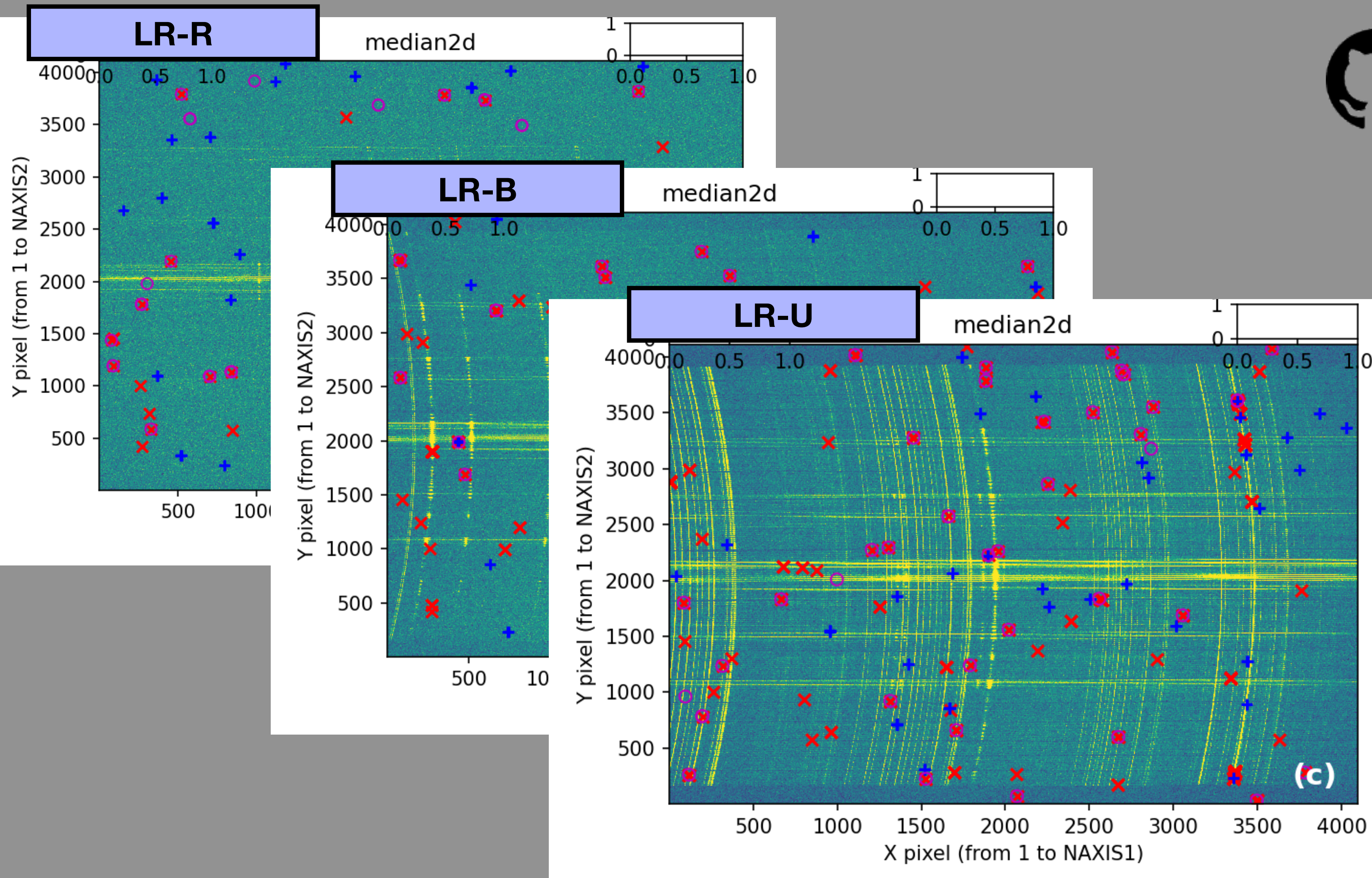
Pixels affected by cosmic rays in more than one exposure (double or triple CRs) have been found to remain uncorrected in the final image.

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MEGARA Data Reduction Pipeline



crmask
guaix-ucm/numina
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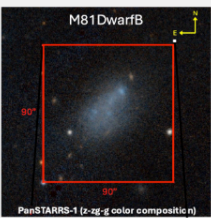
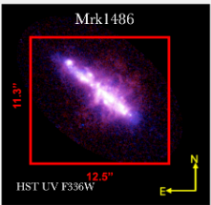
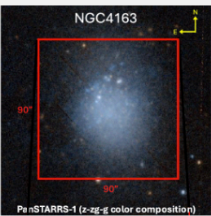
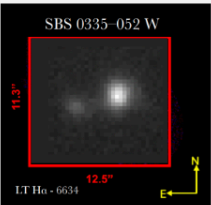
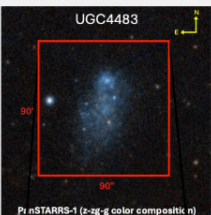
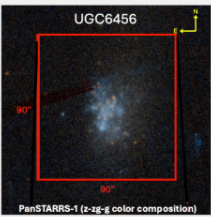


Median combination effectively cleans pixels affected by cosmic rays when they are present in only a single exposure.

Pixels affected by cosmic rays in more than one exposure (double or triple CRs) have been found to remain uncorrected in the final image.

On the order of 100!

What's coming next

<div><div><div><div><div></div><div></div></div><div><div>Dwarfs4</div><div>MOSAIC</div></div></div><div></div></div><div>WELCOME, ADMIN / HOME / DATA</div></div>											
You are here: Home											
<div>Research Objects</div> <div>List of celestial objects intended for study, along with details and relevant properties for each object.</div>											
IMAGE	NAME	TYPE	RA	DEC	MAG	z	SIZE (arcsec)	VISIBILITY SEMESTER	OBSERVING RUN	INSTRUMENT	DATA FILES
	M81DwarfB	galaxy	10:05:30.1	70:21:50.866		0.001147	2.8	2023B2	WS2023B2-023	WEAVE	stackcube_3030307.fits stackcube_3030308.fits
	Mrk1486	galaxy	13:59:50.88	57:26:22.9		0.033829		2023	GTC4-23ITP	MEGARA	final_cube_GTC4-23ITP_OB0001.fits final_cube_GTC4-23ITP_OB0002.fits final_cube_GTC4-23ITP_OB0003.fits
	NGC4163	galaxy	12:12:09.1581	36:10:11.299		0.000547		2023B2	WS2023B2-023	WEAVE	stackcube_3043873.fit stackcube_3043874.fit stackcube_3044407.fit stackcube_3044408.fit
	SBS0335-052W	galaxy	03:37:38.40	-05:02:37.5				2023	GTC4-23ITP	MEGARA	final_cube_GTC4-23ITP_OB0006.fits final_cube_GTC4-23ITP_OB0009.fits
	UGC4483	galaxy	08:37:02.9721	69:46:32.800		0.000519		2023B2	WS2023B2-023	WEAVE	stackcube_3039318.fit stackcube_3039319.fit stackcube_3039505.fit stackcube_3039506.fit
	UGC6456	galaxy	11:27:59.0422	78:59:41.802		-0.000345		2023B2	WS2023B2-023	WEAVE	stackcube_3048085.fit stackcube_3048086.fit stackcube_3048225.fit stackcube_3048226.fit stackcube_3052913.fit stackcube_3052914.fit



Dwarfs4MOSAIC Data Release 1



Study of gas-phase metallicity gradients in dwarf galaxies

(Dwarfs4MOSAIC paper with M.Lara-López as leader “*probing mass assembly in low-mass galaxies through gas metallicity gradients*”)



Reducing new MEGARA images from Dwarfs4MOSAIC sample

Dwarfs4MOSAIC web repository created for Data Release 1 by Noelia Grande Zafrilla (MSc Thesis 2024–2025)

Preliminary Results from Dwarfs4MOSAIC:

A two-dimensional study of low-mass, high star-formation galaxies at low redshift

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